

# The Silicon Vertex Detector of the Belle II Experiment

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The Belle II experiment operating at the asymmetric-energy  $e^+e^-$  SuperKEKB collider, located in Tsukuba (Japan), has been collecting data since March 2019. Its excellent vertexing abilities are provided by the vertex detector, part of which is the silicon-strip vertex detector (SVD) that plays a crucial role in the charged-particle tracking close to the interaction point. The SVD has operated successfully and efficiently over the whole period of data taking so far. In this article, we briefly discuss its purpose, structure and basic description of the front-end electronics. The main quantities related to the SVD performance are presented. The foreseen increase in SuperKEKB luminosity will lead to higher background, so we describe its impact on the SVD performance. A quick overview of the radiation damage campaign is presented to show the predicted behaviour of the sensors subjected to high radiation, whose level is constantly monitored. We also discuss the ongoing software development to account for the high occupancy expected in the future. In particular, the utilization of the SVD hit time information is presented as a very important quantity to suppress off-time background hits and tracks. Finally, the work done during the first long shutdown of SuperKEKB is briefly described, during which a major upgrade of the pixel detector has been successfully done. Resumption of the beam operation is expected in early 2024.

Keywords: Silicon strip detector, Vertex detector, Tracking detector, Belle II

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## 36 1. Introduction

The Belle II [1] experiment is dedicated to search for physics beyond the standard model at 37 the intensity frontier. It operates at the SuperKEKB collider located at KEK, Tsukuba in Japan, 38 providing asymmetric beams of 7 GeV electrons and 4 GeV positrons. In the accelerator's default 39 operation regime, the center-of-mass energy is set to the  $\Upsilon(4S)$  resonance, hence it produces a huge 40 sample of B mesons via the  $e^+e^- \rightarrow \Upsilon(4S) \rightarrow B\bar{B}$  process. So far, SuperKEKB achieved the 41 highest instantaneous luminosity of  $4.7 \times 10^{34}$  cm<sup>-2</sup>s<sup>-1</sup>, which is the current world record. The 42 Belle II detector is a multipurpose spectrometer characterized by excellent vertexing capability and 43 good hermeticity, which has accumulated 424 fb<sup>-1</sup> to date, and its final goal is to collect a data 44 sample of 50  $ab^{-1}$ , that will be possible with a constant increase of the SuperKEKB instantaneous 45 luminosity up to our final goal of  $6 \times 10^{35}$  cm<sup>-2</sup>s<sup>-1</sup>. 46

Belle II is composed of various sub-detectors with the vertex detector (VXD) being the closest 47 to the interaction point. It is divided into two further subsystems. The innermost part is the pixel 48 detector (PXD), which is based on depleted field effect transistor pixel sensors. The PXD consists 49 of two layers (numbered 1-2) and its main goal is the precise determination of the decay vertices. 50 Outside the PXD is the silicon-strip vertex detector (SVD) [2] with four layers (numbered 3-6) 51 that mostly extrapolates the measured tracks to the PXD, defining the so-called region of interest 52 (ROI), which significantly reduces the amount of data recorded by the PXD. The SVD also performs 53 standalone tracking for low-momentum charged particles and contributes to their identification by 54 providing energy loss information. 55

#### 56 2. SVD structure

Each SVD layer is composed of a number of double-sided silicon strip detectors (DSSDs) 57 that are manufactured on an n-type bulk wafer with a thickness of about 300  $\mu$ m (Figure 1). One 58 side of the sensor is covered by the p-type silicon strips placed in parallel to the beam axis that 59 determine the  $r - \phi$  coordinates (distance from the z-axis and azimuthal angle, respectively), and the 60 n-type strips are placed perpendicularly on the other side of the bulk, measuring the z coordinate 61 (collinear to the electron beam). Figure 1 (left) shows a schematic picture of SVD layers and 62 associated sensors with increasing numbering from the forward (FWD) to the backward (BWD) 63 regions. Such structure is repeated along the azimuthal direction forming different ladders and the 64 so-called windmill geometry of the SVD. The sensors differ depending on the layer and the region 65 in which they are placed in the SVD. In the FWD part, for layers 4-6, they have a trapezoidal shape 66



**Figure 1:** Schematic picture of SVD sensors forming different layers (left) and a table summarizing the parameters for each type of sensor (right).

and are slanted in the region that, due to the asymmetric beams, is characterised by the highest track 67

multiplicity. In addition, in layer 3 the sensors are smaller and contain more n-type strips than that 68

in layers 4-6. This also implies the readout pitch (distance between two readout strips) to be much 69

- smaller for p-side strips with respect to the n-side. To improve spatial resolution, a floating strip is 70 placed between two readout strips on both p- and n-sides. The charge induced in the floating strip
- is shared by the neighboring strips, reducing the effective strip pitch to half of the readout pitch. 72
- The right table of Figure 1 summarises the sensor parameters. The SVD consists of 224 thousand 73
- readout strips and 172 sensors with an active area of  $1.2 \text{ m}^2$ . 74

#### 2.1 Fornt-end electronics 75

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For the readout we use APV25 chips [3]. For the central part of SVD (except for layer 3), the 76 chips are attached directly to the DSSD sensors via flex circuits bent over the DSSD edge (Origami 77 concept). The edge sensors use hybrid boards located outside the active volume. The APV25 has 78 128 channels per chip and amplifiers that provide a shaping time of 50 ns. Radiation hardness 79 exceeds 100 Mrad and the power consumption is around 0.4 W/chip. The sampling frequency is 32 80 MHz and after the trigger's arrival we can collect six consecutive signal samples in total with the 81 multipeak mode. To account for higher luminosity in the future, we have introduced the so-called 82 "3/6 mixed acquisition mode", which allows switching between three and six samples recorded on 83 an event-by-event basis, based on the trigger type (and hence its time accuracy) for a particular 84 event. This mode, already prepared and tested, significantly reduces the data size, which can be 85 crucial in high background conditions. 86

#### SVD performance 3. 87

Since the start of the operation we have observed very smooth performance of the SVD, with a 88 very few masked strips (less than 1%). Moreover, the environment has been stable and the evolution 89 of calibration constants is consistent with expectation. Also, the effects of radiation damage are 90 well under control. 91

Several quantities related to the SVD performance - sensor efficiency, signal-to-noise ratio, and 92 both spatial and time resolution - are constantly monitored. Regarding SVD sensor efficiency, the 93 values for all sensors are typically over 99% and they are also very stable over the whole period of 94 data taking. Clusters are formed from adjacent strips with significant signal and the charge collected 95 in a given cluster strongly depends on the incident angle of the track. Over time, we observe very 96 similar cluster charge in all the sensors once normalized to the track's length. For layer 4-5-6 on 97 the n-type strips we observe 10-30% loss of the signal due to the large pitch combined with the 98 presence of a floating strip. Another important quantity is the signal-to-noise ratio (SNR), which is 99 satisfactory for all 172 sensors. The SNR MPV is ranging from 13 to 30, depending on the sensor 100 position, due to the track incident angle with the sensor, and on the sensor side, with smaller SNR 101 for the p-sides, due to larger noise for the longer strip length. A small decrease of cluster SNR value 102 is observed in 2022 measurement, due to increased noise from radiation damage by approximately 103 20%– 30%. In Figure 2 the distributions of cluster charge (left) and SNR (right) are presented, 104 where histograms representing the data collected in 2020 and 2022 are superimposed. 105



**Figure 2:** Distribution of cluster charge (left) and signal-to-noise ratio (right) for layer 3 (p-side). Comparison between data taken in 2020 (blue) and 2022 (red) is presented.

Both position and time resolution are very important metrics for excellent SVD performance. The position resolution measurement is based on the residuals, i.e., the clusters' positions with respect to the intercept of the unbiased tracks' extrapolation, and it is evaluated with a large sample of  $e^+e^- \rightarrow \mu^+\mu^-$  decays. As shown in Figure 3, this quantity depends on the incident angle and is very stable during the period of the Belle II operation. As seen in Figure 3, the resolution for the n-side (left plot) is about two times worse with respect to that for the p-side, which is a result of the different pitch.

Hit time resolution is measured with respect to the event time of the collision provided by 113 central drift chamber (CDC) and exhibits a very good resolution of less then 3 ns for the clusters 114 associated to tracks. Using the average value of all the hits on a given track, the so-called "track-115 time" can be computed, slightly improving the time resolution. Furthermore, the "event-time" can 116 be determined using all the clusters associated to selected tracks in an event. In such a way, the 117 "event time" can be computed by the SVD with a resolution of the order of 1 ns, which is around 118 2000 times faster with respect to the CDC-based computation. This feature will be especially 119 important in the higher luminosity environment, as it can significantly speed up the reconstruction 120 process at the high-level trigger. 121



**Figure 3:** Distributions of position resolution for p-side (left) and n-side (right) as a function of the track incident angle. A comparison between data taken in 2020 (dots) and 2022 (dotted lines) is presented.

#### 122 4. Radiation effects

In the high-energy physics experiments, the effects from radiation damage coming from ma-123 chine related background is a major factor that deteriorates the sensor performance with time. The 124 SVD accumulated dose is constantly measured using data from diamond sensors that are mounted 125 on the IP beam pipe, and the corresponding level of the equivalent neutron fluence is evaluated using 126 the ratio of equivalent neutron fluence to dose estimated from Monte Carlo simulation. Several 127 effects related to radiation damage must be taken into account. A linear increase of the leakage 128 current as a function of radiation damage is observed in the sensors, as expected from the bulk 129 damage described by the NIEL model [4], and shown in Figure 4 left. The sensor current is shown 130 as a function of the accumulated dose for one of the layer 3 sensors most exposed, that received 131 about 70 krad to date, corresponding to about  $1.6 \times 10^{11}$  neq. So far, this increase has had a 132 negligible contribution to the noise because of both the small leakage current and the short APV25 133 shaping time. The rate of the leakage current increase measured is consistent with the experience 134 from other experiments working with similar detectors and in comparable conditions [5]. However, 135 we expect some significant impact on the strip noise due to the sensor leakage current, and hence 136 a deterioration in SNR, for the dose of  $\sim 6$  Mrad, which is thus considered as SVD dose limit 137 to preserve optimal performance. The strip noise for unirradiated modules is dominated by the 138 interstrip capacitance. During the operation we have observed an increase in its value of about 20%139 (30%) for n-side (p-side), due to effects of surface radiation damage that increases the interstrip 140 capacitance, but it is expected to saturate, as also visible in Figure 4 center. 141

Another relevant effect of the bulk radiation damage is the impact on depletion voltage. The expected future radiation levels at the nominal luminosity, of about 0.35 Mrad/year and  $8 \times 10^{11} n_{eq}/cm^2/year$ , are affected by large uncertainty due to the machine evolution as well as a possible redesign of the interaction region. To better explore the possible effects of bulk damage in the SVD sensors after bulk type inversion, an irradiation campaign was conducted in July 2022 at ELPH, Tohoku University. Several SVD sensors have been exposed to a 90 MeV electrons beam,



**Figure 4:** Left plot: Leakage current as a function of the accumulated dose; Center plot: the average noise level as a function of the accumulated dose for the p-side (blue dots) and n-side (red dots); Right plot: full depletion voltage as a function of the accumulated dose with the type inversion observed at 2 Mrad.

up to 10 Mrad, corresponding to an equivalent neutron fluence of  $3 \times 10^{13}$  n<sub>eq</sub>/cm<sup>2</sup>. The decrease 148 of the depletion voltage has been observed up to the point of bulk type inversion, which occurred at 149 2 Mrad (~  $6 \times 10^{12} n_{ea}/cm^2$ ), after which the depletion voltage started to increase again (Figure 4 150 right). Detailed measurements, whose results will be shortly published, confirmed that the sensors 151 will still work fine after the type inversion, which meets our expectation for these types of silicon 152 detectors. Since the beginning of the detector operation, we have not observed any change in the 153 depletion voltage in the sensors installed in the SVD, as expected due to the small accumulated 154 equivalent neutron fluence so far, below  $2 \times 10^{11} n_{eo}/cm^2$ . Considering all these results, the dose 155 limit of 6 Mrad and the extrapolation of the background levels quoted above, the SVD has a wide 156 safety margin for the accumulated radiation damage even after 10 years of the operation at the target 157 luminosity. 158

## 159 5. High background scenario and related software/hardware developments

With the increase of the luminosity and the expected larger machine related background, the 160 SVD occupancy will also increase and a deterioration of the tracking performance is expected above 161 certain levels. So far, the average hit occupancy is 0.5% for layer 3, which does not degrade the 162 performance. Nonetheless, the background extrapolation for different future scenarios has been 163 performed with detailed simulations of the various contributions (beam-gas, Toushek, etc.) and 164 applying appropriate data-simulation scale factors [6]. These studies predict that for the nominal 165 luminosity we can reach an occupancy in layer 3 very close to the limit of 4.7%, above which the 166 tracking performance deteriorates. These predictions have large uncertainties coming from poorly 167 known machine evolution in the future, with a possible redesign of the interaction region. In the 168 most conservative scenario, the layer-3 occupancy can increase up to  $\sim 8.7\%$ , which is far beyond 169 the modest tracking performance. Such a scenario motivates us to develop the SVD reconstruction 170 software, as well as to seriously consider the VXD upgrade [7], since the safety factor might be too 171 small to ensure good quality data. The technology assessment related to this hardware upgrade is 172 currently ongoing. 173

An important effort related to the software development is the utilization of the hit time 174 information from the SVD. The real signal hits come from well-triggered collisions, but the SVD 175 acquisition window ( $\sim 100$  ns) is much wider with respect to the SuperKEKB bunch spacing (6 ns). 176 Therefore, we need to cope with many off-time hits related to the beam-induced background or 177 background from the other bunches. The current selection is based on two requirements: a) time 178 difference between p- and n-side cluster,  $|t_p - t_n| < 20$  ns, and b) the absolute value of the cluster 179 time,  $|t_{p,n}| < 50$  ns. These criteria reject the majority of the background hits retaining above 99% of 180 the signal, and based on them the SVD occupancy limit for layer 3 can be set at 4.7%. Recently, a 181 more effective background suppression method has been developed in the form of so-called "SVD 182 grouping". It is based on an event-by-event classification of the clusters by their time, so the 183 clusters belonging to tracks from the same collisions are collected in the same group. Clusters 184 from the different collisions or beam background will be placed in the other groups; finally, only 185 the clusters belonging to the priority group will be used for the tracking. This feature reduces the 186 fake rate (fraction of the fake tracks) by 16% for the high-background scenario. An additional fake 187 rate reduction can be achieved by applying the selection on the track-time to reject off-time tracks. 188

Finally, these improvements allow an increase of the SVD occupancy limit for layer 3 from 4.7% to around 6%.

#### 191 6. Activities during the Long Shutdown 1

Long shutdown 1 (LS1) started in May 2022 and one of the goals was to upgrade the VXD with a 192 new PXD. During the first data taking period, the second layer of PXD was only partially equipped, 193 and 5/6 of the azimuthal angle remained uncovered. The new PXD provides the full coverage, 194 which is beneficial for more precise vertexing. Hardware activities for the VXD uninstallation and 195 reinstallation were intense: after the VXD extraction from Belle II, the SVD was detached from 196 the old PXD (May 16-17, 2023), then the new PXD was attached to the SVD (June 20-21, 2023) 197 and finally the complete VXD was installed in the Belle II detector. The whole delicate procedure 198 had neither major problems nor caused any damage. In the period of September 12 - October 1, 199 2023, the VXD commissioning was performed to confirm the PXD and SVD performance, and 200 also to check the impact from the increased PXD power consumption and possible increase in 201 the temperature on the sensor leakage current. From September 21, several cosmic runs with no 202 magnetic field were taken to check the performance and compare them with corresponding ones 203 for 2022 data samples. We observed no issues, in particular the noise distributions over readout 204 channels remained basically unchanged as well as SNR for the clusters associated to the tracks, with 205 stable excellent efficiency for all the sensors. 206

#### 207 7. Conclusions

To conclude, SVD has successfully operated since March 2019 with very smooth performance and without major problems. Its good vertexing quality has been confirmed by many physics measurements, in particular those related to the lifetime analyses e.g. Ref. [8]. Some radiation damage effects were observed, but without any impact on the performance so far.

However, the extrapolated background level indicates that the occupancy in the SVD can exceed the current limit that guaranties good tracking performance. Hence, several software improvements are being implemented to account for high background conditions. In particular, exploitation of the SVD hit time is of major importance. Alongside, a VXD upgrade is also under discussion to increase robustness against high background and to match a possible new interaction region.

The VXD reinstallation at Belle II with complete PXD has been successfully done during the LS1, followed by successful VXD commissioning with cosmic data. The beam operation is planned to resume in early 2024.

#### 220 **References**

- [1] T. Abe et al., Belle II Technical Design Report, arXiv:1011.0352 (2010).
- <sup>222</sup> [2] K. Adamczyk et al., JINST **17**, P11042 (2022).
- [3] M. J. French et al., Nucl. Instrum. Meth. A 466, 359 (2001).

- <sup>224</sup> [4] G. Lindstrom et al., Nucl. Instrum. Meth. A **465**, 60-69 (2000).
- <sup>225</sup> [5] B. Aubert et al., Nucl. Instrum. Meth. A **729**, 615 (2013).
- [6] A. Natochii et al., Nucl. Instrum. Meth. A **1055**, 168550 (2023).
- <sup>227</sup> [7] M. Babeluk et. al., Nucl. Instrum. Meth. A **1048**, 168015 (2023).
- <sup>228</sup> [8] F. Abudinén et al., Phys. Rev. Lett. **130**, 071802 (2023)